University of Cologne www.thp.ur Institute for Theoretical Physics Prof. Dr. Claus Kiefer Nick Kwidzinski, Leonardo Chataignier and Yi-Fan Wang

ver. 1.10

## 3<sup>rd</sup> exercise sheet on Relativity and Cosmology II

Summer term 2019

**Release**: Mon, Apr. 15<sup>th</sup>**Submit**: Mon, Apr. 29<sup>th</sup> in lecture**Discuss**: Thu, May 2<sup>nd</sup>Please note that you have two weeks for this exercise sheet. It gives twice the points as usual.

## **Exercise 44** (15 points): *Differential forms*

**44.1** Consider an *n*-dimensional manifold with a metric. Let  $\{\omega^i\}$  be an orthonormal co-basis of 1-forms, and let  $\omega$  be the preferred volume form  $\omega = \omega^1 \wedge \omega^2 \wedge \ldots \wedge \omega^n$ .

Show that, in an arbitrary coordinate system  $\{x^k\}$ , the following holds:

$$\omega = \sqrt{|g|} \, \mathrm{d}x^1 \wedge \mathrm{d}x^2 \wedge \ldots \wedge \mathrm{d}x^n \,, \tag{1}$$

where g denotes the determinant of the metric, whose components  $g_{ij}$  are given in these coordinates.

**44.2** The *contraction* of a *p*-form  $\omega$  (with components  $\omega_{ij\ldots k}$ ) with a vector v (with components  $v^i$ ) is given by  $[\omega(v)]_{j\ldots k} := \omega_{ij\ldots k} v^i$ . Consider the *n*-form  $\omega = dx^1 \wedge dx^2 \wedge \ldots \wedge dx^n$ .

Show that, with a given vector field *v*, the following holds:

$$\mathbf{d}[\omega(v)] = v_{,i}^{i} \,\omega\,. \tag{2}$$

**44.3** Define  $(\operatorname{div}_{\omega} v) \omega := d[\omega(v)]$ .

Show that, by using coordinates in which  $\omega = f dx^1 \wedge dx^2 \wedge \ldots \wedge dx^n$ , the following holds:

$$\operatorname{div}_{\omega} v = \frac{1}{f} \left( f v^{i} \right)_{,i} \,. \tag{3}$$

**44.4** In three-dimensional Euclidean space, the preferred volume form is given by  $\omega = dx \wedge dy \wedge dz$ . Show that, in spherical coordinates, this volume form is given by  $\omega = r^2 \sin \theta \, dr \wedge d\theta \wedge d\phi$ .

Use the result of **44.3** to show that the divergence of a vector field

$$v = v^{r} \frac{\partial}{\partial r} + v^{\theta} \frac{\partial}{\partial \theta} + v^{\phi} \frac{\partial}{\partial \phi}$$
(4)

is given by

div 
$$v = \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 v^r \right) + \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \, v^\theta \right) + \frac{\partial v^\phi}{\partial \phi} \,.$$
 (5)

## **Exercise 45** (25 points): *Electrodynamics in flat space-time*

Differential forms are a convenient tool for field theories, as we will show in this exercise on the example of Maxwell electrodynamics in Minkowski space-time. We know that the electromagnetic field strength is given by the Faraday antisymmetric tensor  $F_{\mu\nu}$ , and the current is given by  $j^{\mu}$ . In the language of exterior calculus, the  $F_{\mu\nu}$  are components of the *Faraday* 2-*form* **F** describing an arbitrary electromagnetic field, given by

$$\mathbf{F} := \frac{1}{2} F_{\mu\nu} \, \mathrm{d}x^{\mu} \wedge \mathrm{d}x^{\nu} = -E_x \, \mathrm{d}t \wedge \mathrm{d}x - E_y \, \mathrm{d}t \wedge \mathrm{d}y - E_z \, \mathrm{d}t \wedge \mathrm{d}z + B_x \, \mathrm{d}y \wedge \mathrm{d}z + B_y \, \mathrm{d}z \wedge \mathrm{d}x + B_z \, \mathrm{d}x \wedge \mathrm{d}y \,, \quad (6)$$

while the  $j^{\mu}$  are the components of the current 1-form **j**, which is given by

$$\mathbf{j} := j_{\mu} \, \mathrm{d}x^{\mu} = -\rho \, \mathrm{d}t + j_{x} \, \mathrm{d}x + j_{y} \, \mathrm{d}y + j_{z} \, \mathrm{d}z \,. \tag{7}$$

See reverse.

**45.1** The Hodge star operator  $\star$  maps *p*-forms to (4 - p)-forms. Therefore, 2-forms are mapped to 2-forms by this operator. The 2-form dual to the Faraday 2-form is the Maxwell 2-form, defined by **G** :=  $\star$ **F**.

The Hodge star operator acts on the 2-form basis as follows

$$\star (\mathrm{d} x^{\mu} \wedge \mathrm{d} x^{\nu}) = \frac{1}{2} \eta^{\mu\alpha} \eta^{\nu\beta} \underbrace{\varepsilon}_{\alpha\beta\gamma\delta} \mathrm{d} x^{\gamma} \wedge \mathrm{d} x^{\delta}, \qquad (8)$$

where  $\eta^{\mu\nu} = \text{diag}(-1, +1, +1, +1)$  is the inverse Minkowski metric and  $\varepsilon_{\alpha\beta\gamma\delta}$  the totally anti-symmetric Levi-Civita tensor density ( $\varepsilon_{0123} = +1$ ). For example,

$$\star (\mathrm{d}t \wedge \mathrm{d}x) = -\mathrm{d}y \wedge \mathrm{d}z, \quad \text{etc.} \tag{9}$$

Show that the Maxwell 2-form is given by

$$\mathbf{G} = B_x \, \mathrm{d}t \wedge \mathrm{d}x + B_y \, \mathrm{d}t \wedge \mathrm{d}y + B_z \, \mathrm{d}t \wedge \mathrm{d}z + E_x \, \mathrm{d}y \wedge \mathrm{d}z + E_y \, \mathrm{d}z \wedge \mathrm{d}x + E_z \, \mathrm{d}x \wedge \mathrm{d}y \,. \tag{10}$$

Which formal relation between the components of F and G holds?

- **45.2** With these definitions, the Maxwell equations (in Gaussian units) can be written in a compact form:  $d\mathbf{F} = 0$  and  $d\mathbf{G} = 4\pi \star \mathbf{j}$ .
  - a) Show that the equation  $d\mathbf{F} = 0$  corresponds to the two homogeneous Maxwell equations

$$\vec{\nabla} \cdot \vec{B} = 0 \quad \text{and} \quad \vec{\nabla} \times \vec{E} + \partial_t \vec{B} = 0.$$
 (11)

b) Calculate **\*j** (a 3-form), which is the dual of the current 1-form **j**. For that purpose use the following relations

$$\star \mathrm{d}x^{\mu} = \frac{1}{3!} \eta^{\mu\alpha} \underline{\varepsilon}_{\alpha\beta\gamma\delta} \,\mathrm{d}x^{\beta} \wedge \mathrm{d}x^{\gamma} \wedge \mathrm{d}x^{\delta} \,, \tag{12}$$

for example

$$\star dt = -dx \wedge dy \wedge dz, \quad \text{etc.} \tag{13}$$

c) Show that the equation  $d\mathbf{G} = 4\pi \star \mathbf{j}$  corresponds to the two inhomogeneous Maxwell equations

$$\vec{\nabla} \cdot \vec{E} = 4\pi\rho \quad \text{and} \quad \vec{\nabla} \times \vec{B} - \partial_t \vec{E} = 4\pi \vec{j}.$$
 (14)

- d) Which formal manipulations can you perform to turn eq. (11) into eq. (14) in the vacuum case?
- **45.3** The exterior derivative is nilpotent, i.e. the relation  $d(d\omega) = 0$  holds for any *p*-form  $\omega$ .

Choosing  $\omega = \star \mathbf{j}$ , show that this yields the continuity equation  $\partial_t \rho - \vec{\nabla} \cdot \vec{j} = 0$ .

45.5 Derive the inhomogeneous Maxwell equations from a variational principle. The action is given by

$$S[\mathbf{A}] = \int \left(\frac{1}{8\pi} \star \mathbf{F} \wedge \mathbf{F} + \mathbf{A} \wedge \star \mathbf{j}\right) , \qquad (15)$$

where  $\mathbf{F} = d\mathbf{A}$  and  $\mathbf{A} = A_{\mu} dx^{\mu}$  is the electromagnetic 1-form potential.

Show in addition that the continuity equation implies that the action is invariant under the gauge transformation  $\mathbf{A} \rightarrow \mathbf{A}' = \mathbf{A} + df$ , where *f* is a 0-form.

*Hint*: You can make use of the fact that  $\star \alpha \land \beta = \alpha \land \star \beta$ , where  $\alpha$  and  $\beta$  are both *k*-forms.

**45.6** *Bonus exercise.* Let  $\omega$  be a *k*-form. We define the *codifferential*  $\delta$  via

$$\delta \boldsymbol{\omega} := (-1)^k \star \mathbf{d} \star \boldsymbol{\omega} \,. \tag{16}$$

This allows us to define the d'Alembert operator

$$\Box := \mathrm{d}\delta + \delta\mathrm{d} \;. \tag{17}$$

Show the following statements:

- $\delta A = 0$  is the well-known Lorenz gauge condition  $\partial_{\mu} A^{\mu} = 0$
- The Lorenz gauge can always be realized by a suitable gauge transformation.
- In the Lorenz gauge the inhomogeneous Maxwell equation can be written as  $\Box A = 4\pi \mathbf{j}$ .
- **45.7** *Bonus exercise.* Explain in your own words why a covariant derivative is needed on a curved background, and give special attention to the connection. What is the intuitive interpretation of the connection? Why can it be chosen to be zero in Minkowski space-time?