## $4^{\text {th }}$ exercise sheet on Relativity and Cosmology II

Summer term 2019

Release: Mon, Apr. $29^{\text {th }}$
Submit: Mon, May $6^{\text {th }}$ in lecture
Discuss: Thu, May $9^{\text {th }}$

## Exercise 46 (20 credit points): Derivation of the Schwarzschild solution in Cartan calculus

The aim of this exercise is to derive the Schwarzschild solution using the Cartan formalism.
The general spherically symmetric ansatz is given by:

$$
\begin{equation*}
\mathrm{d} s^{2}=g_{\mu \nu} \mathrm{d} x^{\mu} \otimes \mathrm{d} x^{\nu}=-\mathrm{e}^{2 a(r, t)} \mathrm{d} t^{2}+\mathrm{e}^{2 b(r, t)} \mathrm{d} r^{2}+r^{2} \mathrm{~d} \theta^{2}+r^{2} \sin ^{2} \theta \mathrm{~d} \phi^{2} . \tag{1}
\end{equation*}
$$

In terms of the pseudo-orthogonal coframe basis $\left\{\vartheta^{i}\right\}, i=0, \ldots, 3$, the metric takes the form

$$
\begin{equation*}
\mathrm{d} s^{2}=\eta_{i j} \vartheta^{i} \otimes \vartheta^{j}=-\vartheta^{0} \otimes \vartheta^{0}+\vartheta^{1} \otimes \vartheta^{1}+\vartheta^{2} \otimes \vartheta^{2}+\vartheta^{3} \otimes \vartheta^{3} . \tag{2}
\end{equation*}
$$

In this exercise, numerals $(0,1,2,3)$ and Latin indices $(i, j, k, l, \ldots)$ indicate an anholonomic basis, whereas coordinate $(t, r, \theta, \phi)$ and Greek indices $(\mu, v, \lambda, \sigma, \ldots)$ refer to a holonomic basis. Holonomic and anholonomic indices can be converted into each other by using the tetrad defined via

$$
\begin{equation*}
\vartheta^{i}=e_{\mu}{ }^{i} \mathrm{~d} x^{\mu} \tag{3}
\end{equation*}
$$

46.1 Show that

$$
\begin{equation*}
\vartheta^{0}=\mathrm{e}^{a(r, t)} \mathrm{d} t, \quad \vartheta^{1}=\mathrm{e}^{b(r, t)} \mathrm{d} r, \quad \vartheta^{2}=r \mathrm{~d} \theta, \quad \vartheta^{3}=r \sin \theta \mathrm{~d} \phi \tag{4}
\end{equation*}
$$

defines a suitable orthonormal coframe basis to describe the spherically symmetric ansatz.
46.2 Now, calculate the exterior derivatives $\mathrm{d} \vartheta^{i}$. Insert these into the first Cartan structure equation

$$
\begin{equation*}
\Theta^{i}:=\mathrm{d} \vartheta^{i}+\omega^{i}{ }_{j} \wedge \vartheta^{j}=0 \tag{5}
\end{equation*}
$$

and show that the connection $\omega^{i}{ }_{j}$ in components reads

$$
\begin{array}{lll}
\omega^{1}{ }_{0}=a^{\prime} \mathrm{e}^{-b(t, r)} \vartheta^{0}+\dot{b} \mathrm{e}^{-a(t, r)} \vartheta^{1}, & \omega^{3}{ }_{1}=\frac{\mathrm{e}^{-b(t, r)}}{r} \vartheta^{3}, & \omega^{3}{ }_{2}=\frac{\cot \theta}{r} \vartheta^{3}, \\
\omega^{2}{ }_{1}=\frac{\mathrm{e}^{-b(t, r)}}{r} \vartheta^{2}, & \omega^{3}{ }_{0}=0, & \omega^{2}{ }_{0}=0,
\end{array}
$$

where a prime and superscript dot denote derivatives with respect to $r$ and $t$, respectively.
Hint: For any metric-compatible connection (like the Levi-Civita connection studied here) we have $\nabla g_{i j}=0$ and therefore $\omega_{i j}=-\omega_{j i}$.
46.3 Calculate the curvature 2 -forms $\Omega^{i}{ }_{j}:=\mathrm{d} \omega^{i}{ }_{j}+\omega^{i}{ }_{a} \wedge \omega^{a}{ }_{j}$. Use the second Cartan structure equation

$$
\begin{equation*}
\Omega^{i}{ }_{j}=\frac{1}{2} R_{j k l}^{i} \vartheta^{k} \wedge \vartheta^{l} \tag{6}
\end{equation*}
$$

to show that the non-vanishing anholonomic components $R^{i}{ }_{j k l}$ of the Riemann curvature tensor are

$$
\begin{aligned}
-R^{0}{ }_{101} & =\mathrm{e}^{-2 b}\left(a^{\prime 2}-a^{\prime} b^{\prime}+a^{\prime \prime}\right)+\mathrm{e}^{-2 a}\left(\dot{a} \dot{b}-\dot{b}^{2}-\ddot{b}\right)=R^{1}{ }_{010}, \\
R^{0}{ }_{202} & =R^{0}{ }_{303}=-\frac{a^{\prime} \mathrm{e}^{-2 b}}{r}, \quad R^{1}{ }_{212}=R^{1}{ }_{313}=\frac{b^{\prime} \mathrm{e}^{-2 b}}{r}, \\
R^{0}{ }_{212} & =R^{0}{ }_{313}=-\frac{\dot{b} \mathrm{e}^{-a-b}}{r}=-R^{1}{ }_{202}=-R^{1}{ }_{303}, \quad R^{3}{ }_{232}=\frac{1-\mathrm{e}^{-2 b}}{r^{2}} .
\end{aligned}
$$

46.4 Determine the anholonomic components of the Ricci tensor $R_{i j}=R_{i k j}^{l}$ as well as the Ricci scalar $R=\eta^{i j} R_{i j}$. Note that for the contraction of anholonomic indices the Minkowski metric has to be used.
46.5 For a diagonal metric the holonomic components of a (1,1)-tensor coincide with the anholonomic components. Therefore calculate the mixed components of the Einstein tensor

$$
\begin{equation*}
G^{i}{ }_{j} \stackrel{*}{=} G^{\mu}{ }_{v}=R^{\mu}{ }_{v}-\frac{1}{2} \delta_{v}^{u} R . \tag{7}
\end{equation*}
$$

46.6 Back to physics. Outside of the mass distribution we have a vacuum and therefore $T^{\mu}{ }_{v}=0$, i.e. $G^{\mu}{ }_{v}=0$. Use this to show that $b$ depends only on $r$ and that $a(r)=-b(r)$.
46.7 Finally, integrate the differential equation arising from $G_{t}^{t}=0$ to find a relation between $b$ and $r$ and use the Newtonian limit to fix the integration constant. Show that

$$
\begin{equation*}
\mathrm{e}^{-2 b(r)}=1-\frac{2 G M}{r} \tag{8}
\end{equation*}
$$

which concludes our derivation of the Schwarzschild solution.

