ver. 1.00

## 8<sup>th</sup> exercise sheet on Relativity and Cosmology II

Summer term 2019

**Exercise 54** (20 credit points): *Kerr–Newman metric* 

The most general solution for a stationary black hole is given by the *Kerr–Newman metric*, which describes a black hole with angular momentum J = Ma and charge q. The line element expressed in Boyer–Lindquist coordinates takes the following form:

$$\mathrm{d}s^{2} = -\frac{\Delta}{\rho^{2}}\left(\mathrm{d}t - a\sin^{2}(\theta)\,\mathrm{d}\phi\right)^{2} + \frac{\sin^{2}(\theta)}{\rho^{2}}\left[\left(r^{2} + a^{2}\right)\mathrm{d}\phi - a\,\mathrm{d}t\right]^{2} + \frac{\rho^{2}}{\Delta}\,\mathrm{d}r^{2} + \rho^{2}\,\mathrm{d}\theta^{2}\,,$$

where

$$\rho^2 = r^2 + a^2 \cos^2(\theta)$$
,  $\Delta = r^2 - 2Mr + q^2 + a^2$ ,  $q^2 + a^2 \le M^2$ .

- **54.1** How can one obtain this line element from the Kerr metric by means of a simple substitution of certain parameters (without calculation)?
- **54.2** For  $\Delta = 0$  the metric exhibits coordinate singularities. Determine their radial coordinates  $r_{\pm}$ .

The surface  $r_+ = \text{const.}$  (with  $r_+$  being the radial coordinate with a larger value) represents the event horizon. Calculate its surface area for t = const.

**54.3** Analogously to the Kerr metric, consider an observer with r = const.,  $\theta = \pi/2$ , whose tangent vector is parallel to the Killing field  $\chi^{\mu} = \xi^{\mu} + \Omega \Psi^{\mu}$ .

Which values can  $\Omega$  take for given  $r \ge r_+$ ? Show that at the horizon only one value  $\Omega_H$  is possible and determine this value.

**54.4** Consider the Killing field  $\chi^{\mu} = \xi^{\mu} + \Omega \Psi^{\mu}$  evaluated at the event horizon.

Show that this Killing field is light-like on the entire horizon. Furthermore, show that the surface gravity  $\kappa$  defined by means of  $\left[\nabla^{\mu}(\chi_{\nu}\chi^{\nu})\right]_{H} = -2\kappa\chi^{\mu}|_{H}$  is a well-defined quantity.

Calculate the Lie derivative of the defining equation for  $\kappa$  with respect to  $\chi^{\mu}$  and thereby show that  $\kappa$  is constant along the integral curves of  $\chi$ .

*Remark:* After a rather long calculation one obtains  $\kappa = (r_+ - M)/(r_+^2 + a^2)$ . (Not to be shown here.)

**54.5** Consider the null geodesics defined at the horizon, whose tangent vectors  $k^{\mu}$  are proportional to  $\chi^{\mu}$ . Find the functional relationship between the affine parameter  $\lambda$  of these null geodesics and the Killing parameter v of the integral curves of  $\chi^{\mu}$ , i.e.  $\chi^{\mu} = (\partial/\partial v)^{\mu}$ .