

6th exercise sheet on Relativity and Cosmology I

Winter term 2012/13

Deadline for delivery: Thursday, 22nd November 2012 during the exercise class.

Exercise 16 (5 credit points): *Curvature II*

Consider a family of Gaussian curves $z = \exp(-a^2 r^2)$ with $r^2 = x^2 + y^2$, embedded into a flat 3-dimensional space. Determine the metric on a Gaussian curve using polar coordinates (r, φ) and calculate the curvature at the apex using two different methods:

- 16.1** Use the two formulae given in the lecture (comparison of circumference and comparison of area).
16.2 Approximate a spherical shell and use the known curvature of a sphere with radius R .

Exercise 17 (10 credit points): *On the covariant derivative*

- 17.1** In order to define the general covariant derivative $\tilde{\nabla}_\mu$, one only needs a connection $\tilde{\Gamma}^\nu_{\mu\lambda}$, where at first the connection is not linked to the metric in any way:

$$\tilde{\nabla}_\mu T^\nu_\kappa = \partial_\mu T^\nu_\kappa + \tilde{\Gamma}^\nu_{\mu\lambda} T^\lambda_\kappa - \tilde{\Gamma}^\lambda_{\mu\kappa} T^\nu_\lambda.$$

Show that the two conditions

$$Q_{\mu\nu\kappa} \equiv -\tilde{\nabla}_\mu g_{\nu\kappa} = 0, \quad T^\mu_{\nu\kappa} \equiv 2\tilde{\Gamma}^\mu_{[\nu\kappa]} = 0$$

are equivalent to the condition that the connection $\tilde{\Gamma}^\mu_{\nu\kappa}$ is the Christoffel symbol of second kind $\Gamma^\mu_{\nu\kappa}$.

The quantities $Q_{\mu\nu\kappa}$ und $T^\mu_{\nu\kappa}$ are called *non-metricity* and *torsion*, respectively. Are they tensors?

- 17.2** Let V^μ be a vector field and $\mathcal{V}^\mu \equiv \sqrt{-g} V^\mu$ be the corresponding vector density. Show that

$$\nabla_\mu V^\mu = \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} V^\mu) \quad \text{and} \quad \nabla_\mu \mathcal{V}^\mu = \partial_\mu \mathcal{V}^\mu.$$

- 17.3** The covariant wave operator for a scalar field ϕ is given by

$$\square\phi \equiv \nabla^\mu \nabla_\mu \phi.$$

Rewrite this by means of **17.2** such that the resulting expression only contains partial derivatives.

As an example, calculate the wave operator in 3-dimensional spherical coordinates.

Exercise 18 (5 credit points): *Riemannian normal coordinates*

Show that near the origin of a Riemannian normal coordinate system ($\xi^\mu \ll 1$) the following holds:

$$g_{\mu\nu}(0 + \xi) = \eta_{\mu\nu} + \frac{1}{3} R_{\mu\kappa\lambda\nu}(0) \xi^\kappa \xi^\lambda + \dots$$

Give a physical interpretation.

Exercise B₁ (8 bonus points): *Stereographic projection*

Consider a unit 2-sphere.

- B_{1.1}** Introduce stereographic coordinates (x_N, y_N) and (x_S, y_S) by projecting the north and south pole, respectively, onto the equatorial plane and express these coordinates in terms of 3-dimensional cartesian coordinates (x, y, z) . What is the domain of definition of these coordinate mappings?
- B_{1.2}** Determine the coordinates (x, y, z) on the sphere as a function of (x_N, y_N) and (x_S, y_S) , respectively, and thereby calculate the 2-dimensional metric which is induced on the sphere by the flat metric $ds^2 = dx^2 + dy^2 + dz^2$ in terms of both coordinate systems.
- B_{1.3}** Determine explicitly the transfer function on the overlap of the domains of definition and show that the unit sphere combined with both coordinate mappings forms a differentiable manifold.
- B_{1.4}** Verify the transformation formula for the metric by inserting the expressions obtained above:

$$g_{jl}^S = \frac{\partial x_N^i}{\partial x_S^j} \frac{\partial x_N^k}{\partial x_S^l} g_{ik}^N.$$