

## 8<sup>th</sup> exercise sheet on Relativity and Cosmology II

Summer term 2018

**Deadline for delivery:** Thursday, 21<sup>st</sup> June 2018 during the exercise class.

### Exercise 44 (16 credit points + 4 bonus): *Kerr–Newman metric*

The most general solution for a stationary black hole is given by the *Kerr–Newman metric*, which describes a black hole with angular momentum  $J = M a$  and charge  $q$ . The line element expressed in Boyer–Lindquist coordinates takes the following form:

$$ds^2 = -\frac{\Delta}{\rho^2} (dt - a \sin^2(\theta) d\phi)^2 + \frac{\sin^2(\theta)}{\rho^2} [(r^2 + a^2) d\phi - a dt]^2 + \frac{\rho^2}{\Delta} dr^2 + \rho^2 d\theta^2,$$

where

$$\rho^2 = r^2 + a^2 \cos^2(\theta), \quad \Delta = r^2 - 2Mr + q^2 + a^2, \quad q^2 + a^2 \leq M^2.$$

**44.1** How can one obtain this line element from the Kerr metric by means of a simple substitution of certain parameters (without calculation)?

**44.2** For  $\Delta = 0$  the metric exhibits coordinate singularities. Determine their radial coordinates  $r_{\pm}$ .

The surface  $r_+ = \text{const.}$  (with  $r_+$  being the radial coordinate with a larger value) represents the event horizon. Calculate its surface area for  $t = \text{const.}$

**44.3** Analogously to the Kerr metric, consider an observer with  $r = \text{const.}$ ,  $\theta = \pi/2$ , whose tangent vector is parallel to the Killing field  $\chi^\mu = \xi^\mu + \Omega \Psi^\mu$ .

Which values can  $\Omega$  take for given  $r \geq r_+$ ? Show that at the horizon only one value  $\Omega_H$  is possible and determine this value.

**44.4** Consider the Killing field  $\chi^\mu = \xi^\mu + \Omega \Psi^\mu$  evaluated at the event horizon.

Show that this Killing field is light-like on the entire horizon. Furthermore, show that the surface gravity  $\kappa$  defined by means of  $[\nabla^\mu(\chi_\nu \chi^\nu)]_H = -2\kappa \chi^\mu|_H$  is a well-defined quantity.

Calculate the Lie derivative of the defining equation for  $\kappa$  with respect to  $\chi^\mu$  and thereby show that  $\kappa$  is constant along the integral curves of  $\chi$ .

*Remark:* After a rather long calculation one obtains  $\kappa = (r_+ - M)/(r_+^2 + a^2)$ . (Not to be shown here.)

**44.5 Bonus.** Consider the null geodesics defined at the horizon, whose tangent vectors  $k^\mu$  are proportional to  $\chi^\mu$ . Find the functional relationship between the affine parameter  $\lambda$  of these null geodesics and the Killing parameter  $v$  of the integral curves of  $\chi^\mu$ , i.e.  $\chi^\mu = (\partial/\partial v)^\mu$ .

### Exercise 45 (4 credit points): *Hawking temperature*

In the lecture it was mentioned that a Schwarzschild black hole radiates with the so-called *Hawking temperature*

$$T_H = \frac{\hbar c^3}{8\pi k_B G M}.$$

Assume that only photons are emitted and that they have a perfect Planck spectrum. Find a relation between the initial mass of the black hole and its lifetime and analyze this relation for several interesting masses and time intervals.