

## Fourth exercise sheet on Relativity and Cosmology II

### Summer term 2021

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#### Exercise 43 (15 credit points): *Differential forms*

**43.1** Consider a  $n$ -dimensional manifold with a metric. Let  $\{\omega^i\}$  be an orthonormal co-basis of 1-forms, and let  $\omega$  be the preferred volume form  $\omega = \omega^1 \wedge \omega^2 \wedge \dots \wedge \omega^n$ . Show that, in an arbitrary coordinate system  $\{x^k\}$ , the following holds:

$$\omega = \sqrt{|g|} dx^1 \wedge dx^2 \wedge \dots \wedge dx^n, \quad (1)$$

where  $g$  denotes the determinant of the metric, whose components  $g_{ij}$  are given in these coordinates.

**43.2** The contraction of a  $p$ -form  $\omega$  (with components  $\omega_{ij\dots k}$ ) with a vector  $v$  (with components  $v^i$ ) is given by  $[\omega(v)]_{j\dots k} = \omega_{ij\dots k} v^i$ . Consider the  $n$ -form  $\omega = dx^1 \wedge dx^2 \wedge \dots \wedge dx^n$ .

Show that, with a given vector field  $v$ , the following holds:

$$d[\omega(v)] = v^i{}_{,i} \omega. \quad (2)$$

**43.3** Define  $(\operatorname{div}_\omega v) \omega := d[\omega(v)]$ .

Show that, by using coordinates in which  $\omega$  has the form  $\omega = f dx^1 \wedge dx^2 \wedge \dots \wedge dx^n$ , the following holds:

$$\operatorname{div}_\omega v = \frac{1}{f} (f v^i)_{,i}. \quad (3)$$

**43.4** In three-dimensional Euclidean space, the preferred volume form is given by  $\omega = dx \wedge dy \wedge dz$ .

Show that, in spherical coordinates, this volume form is given by  $\omega = r^2 \sin \theta dr \wedge d\theta \wedge d\phi$ .

Use the result of 43.3 to show that the divergence of a vector field

$$v = v^r \frac{\partial}{\partial r} + v^\theta \frac{\partial}{\partial \theta} + v^\phi \frac{\partial}{\partial \phi}, \quad (4)$$

is given by

$$\operatorname{div} v = \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 v^r) + \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} (\sin \theta v^\theta) + \frac{\partial v^\phi}{\partial \phi}. \quad (5)$$

#### Exercise 44 (5 credit points + 8 bonus points): *Electrodynamics in flat space-time*

Differential forms are a convenient tool for field theories, as we will show in this exercise on the example of Maxwell electrodynamics in Minkowski space-time. We know that electromagnetic field strength is given by Faraday antisymmetric tensor  $F_{\mu\nu}$ , and the current is given by  $j^\mu$ . In the language of exterior calculus,  $F_{\mu\nu}$  are components of the Faraday 2-form  $\mathbf{F}$  describing an arbitrary electromagnetic field, given by

$$\mathbf{F} := \frac{1}{2} F_{\mu\nu} dx^\mu \wedge dx^\nu = -E_x dt \wedge dx - E_y dt \wedge dy - E_z dt \wedge dz + B_x dy \wedge dz + B_y dz \wedge dx + B_z dx \wedge dy, \quad (6)$$

while  $j^\mu$  are components of the current 1-form  $\mathbf{j}$  is given by

$$\mathbf{j} := j_\mu dx^\mu = -\rho dt + j_x dx + j_y dy + j_z dz. \quad (7)$$

**44.1** The Hodge star operator  $\star$  maps  $p$ -forms to  $(4-p)$ -forms. Therefore, 2-forms are mapped to 2-forms by this operator. The 2-form dual to the Faraday 2-form is the Maxwell 2-form, defined by  $\mathbf{G} := \star\mathbf{F}$ .

The Hodge star operator acts on the 2-form basis as follows

$$\star(dx^\mu \wedge dx^\nu) = \frac{1}{2} \eta^{\mu\alpha} \eta^{\nu\beta} \varepsilon_{\alpha\beta\gamma\delta} dx^\gamma \wedge dx^\delta, \quad (8)$$

where  $\eta^{\mu\nu} = \text{diag}(-1, +1, +1, +1)$  is the inverse Minkowski metric and  $\varepsilon_{\alpha\beta\gamma\delta}$  totally anti-symmetric Levi-Civita tensor density ( $\varepsilon_{0123} = +1$ ). For example,

$$\star(dt \wedge dx) = -dy \wedge dz, \quad \text{etc.} \quad (9)$$

Show that the Maxwell 2-form is given by

$$\mathbf{G} = B_x dt \wedge dx + B_y dt \wedge dy + B_z dt \wedge dz + E_x dy \wedge dz + E_y dz \wedge dx + E_z dx \wedge dy. \quad (10)$$

Which relation between the components of  $\mathbf{F}$  and  $\mathbf{G}$  holds?

**44.2 Bonus exercise.** With these definitions, the Maxwell equations can be written in a compact form:  $d\mathbf{F} = 0$  and  $d\mathbf{G} = 4\pi \star \mathbf{j}$ , therefore,

a) Show that the equation  $d\mathbf{F} = 0$  corresponds to the two homogeneous Maxwell equations

$$\vec{\nabla} \cdot \vec{B} = 0 \quad \text{and} \quad \vec{\nabla} \times \vec{E} + \partial_t \vec{B} = 0. \quad (11)$$

b) Calculate  $\star \mathbf{j}$  (a 3-form), which is the dual of the current 1-form  $\mathbf{j}$ . For that, use the following relations

$$\star dx^\mu = \frac{1}{3!} \eta^{\mu\alpha} \varepsilon_{\alpha\beta\gamma\delta} dx^\beta \wedge dx^\gamma \wedge dx^\delta, \quad (12)$$

for example

$$\star dt = -dx \wedge dy \wedge dz, \quad \text{etc.} \quad (13)$$

c) Then show that the equation  $d\mathbf{G} = 4\pi \star \mathbf{j}$  corresponds to the two inhomogeneous Maxwell equations

$$\vec{\nabla} \cdot \vec{E} = 4\pi\rho \quad \text{and} \quad \vec{\nabla} \times \vec{B} - \partial_t \vec{E} = 4\pi\vec{j}. \quad (14)$$

d) What can you do to turn ?? into ?? in the vacuum case?

**44.3 Bonus exercise.** The exterior derivative is nilpotent, i.e. the relation  $d(d\omega) = 0$  holds for any  $p$ -form  $\omega$ . Choosing  $\omega = \star \mathbf{j}$ , show that this yields the continuity equation  $\partial_t \rho - \vec{\nabla} \cdot \vec{j} = 0$ .

**44.4 Bonus exercise.** Explain in your own words why a covariant derivative is needed on a curved background, and give special attention to the connection. What is the intuitive interpretation of the connection? Why can it be chosen to be zero in Minkowski space-time?