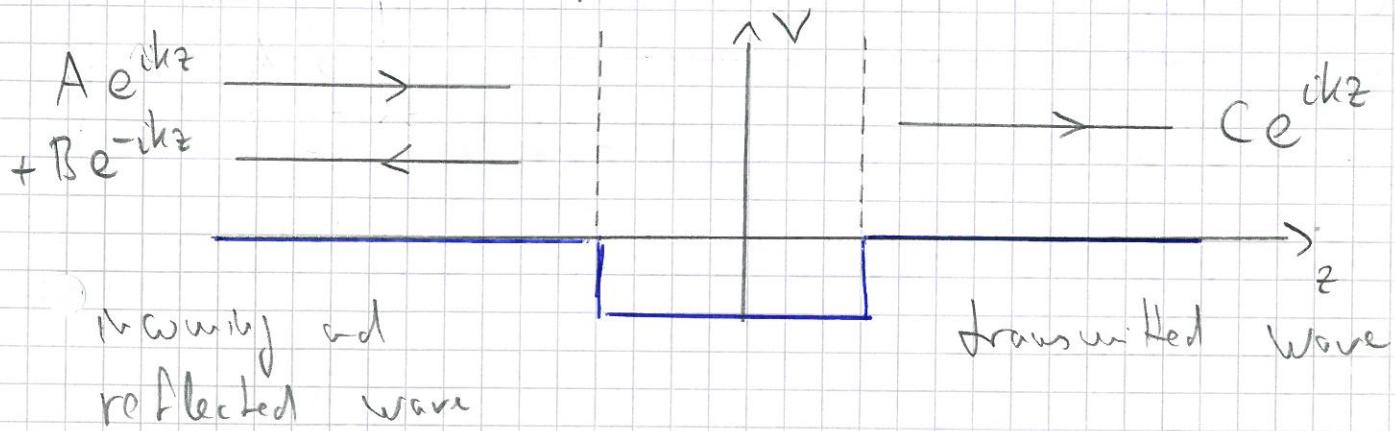


## VI. Scattering theory

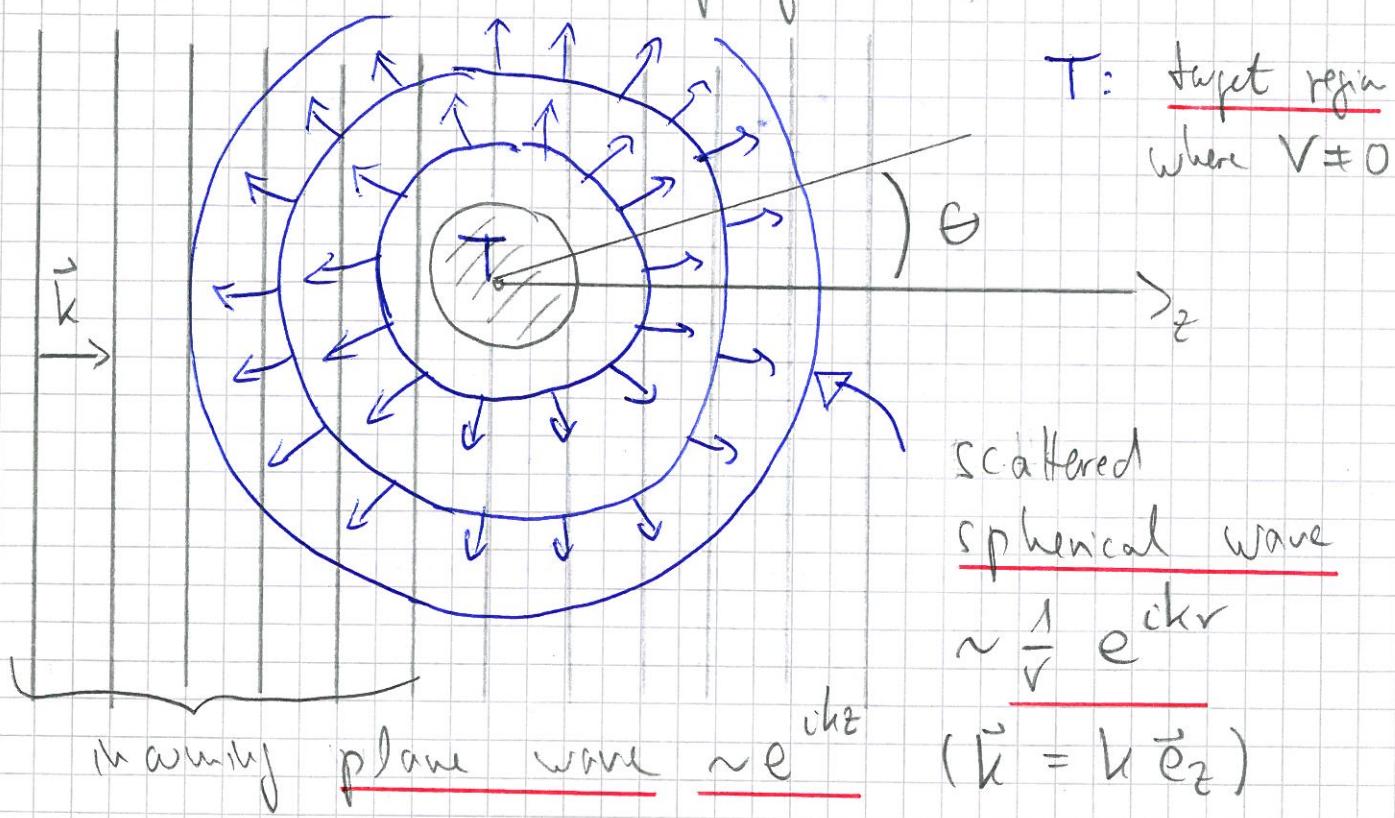
### 1<sup>o</sup> Stationary scattering (elastic throughout)

Reminder: Scattering is one dimension ( $\rightarrow$  Problem 12.1)



Wave functions outside the potential region are solutions of the stationary Schrödinger equation with energy  $E = \frac{\hbar^2 k^2}{2m} > 0$  and momentum  $\vec{k}$ .

### Three-dimensional scattering geometry:



⇒ outside of the target region we expect a solution of the stationary Schrödinger equation with the asymptotic form ( $r \rightarrow \infty$ )

$$\textcircled{S} \quad \psi(r) = A \left( e^{ikz} + f(\theta, \phi) \frac{e^{ikr}}{r} \right) \quad \boxed{\quad}$$

$f(\theta, \phi)$  : scattering amplitude

The probability current associated with this wave function is

$$\vec{j} = -i \frac{e}{2m} [\psi^* \nabla \psi - \psi \nabla \psi^*] \quad \boxed{\quad}$$

Specifically, the incoming current is

$$\vec{j}_{in} = |A|^2 \frac{e k}{m} \hat{e}_z$$

and the radial component of the scattered current

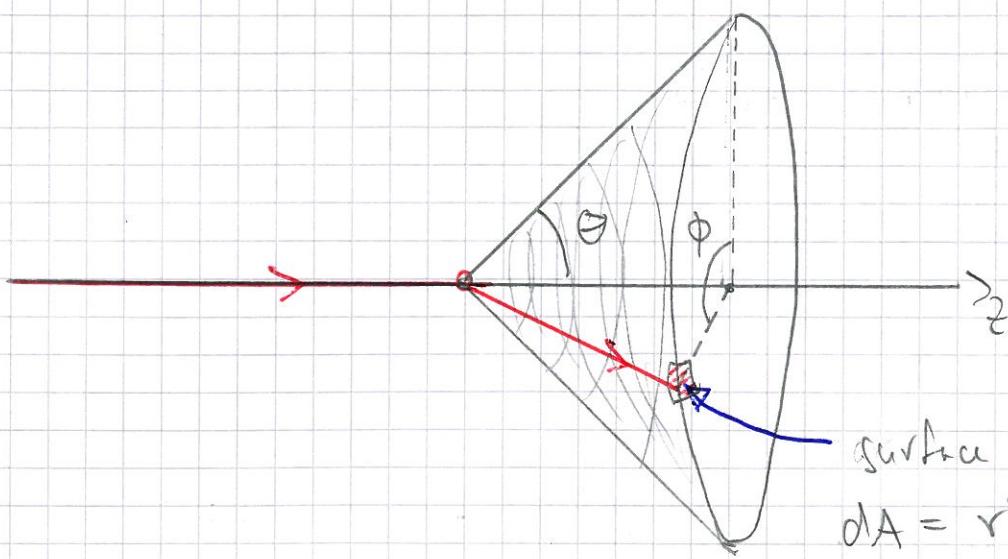
$$j_s^r = -i \frac{e}{2m} \left[ \psi_s^* \frac{\partial \psi_s}{\partial r} - \psi_s \frac{\partial \psi_s^*}{\partial r} \right]$$

with  $\psi_s = A f(\theta, \phi) \frac{e^{ikr}}{r}$ . Thus we have

$$\boxed{j_s^r = |A|^2 |f(\theta, \phi)|^2 \frac{e k}{m} \frac{1}{r^2}}$$

The other contributions to  $j_s^r$  are negligible for large  $r$ .

## Scattering cross section :



$$\begin{aligned} \text{surface element} \\ dA = r^2 d\Omega = \\ = r^2 \sin \theta d\theta d\phi \end{aligned}$$

The differential scattering cross section is the defined as follows:

$$\underline{d\sigma} := \frac{\text{Number of particles through } dA \text{ per time}}{\text{incident current}}$$

$$= \frac{r^2 d\Omega j_{\text{scatt}}}{j_{\text{in}}} = \underline{|f(\theta, \phi)|^2 d\Omega}$$

$$\Rightarrow \underline{\frac{d\sigma}{d\Omega}} = |f(\theta, \phi)|^2$$

Note that it follows from the definition of  $f(\theta, \phi)$  in (5) that  $|f(\theta, \phi)|^2$  has units of an area.

The total scattering cross section is the

$$\underline{\sigma_{\text{tot}}} = \int d\Omega \frac{d\sigma}{d\Omega} = \int d\phi \int d\theta \sin \theta |f(\theta, \phi)|^2 =$$

(175)

$$= \frac{j_{\text{tot}}}{j_{\text{in}}} \quad \text{where the integrated scattered}$$

current is

$$j_{\text{tot}} = \int_0^{2\pi} d\phi \int_0^\pi d\theta r^2 \sin\theta j_{\text{scatt}} =$$

$$= |A|^2 \frac{\pi k}{\omega} b_{\text{tot}}.$$


---

Problems to be addressed:

- A. Find solutions of the stationary Schrödinger equation that satisfy the asymptotic boundary condition  $\textcircled{S}$
  - B. Compute  $f(\theta, \phi)$  for a given potential
  - C. Relate the stationary approach to a more realistic (time-dependent) picture of the scattering process.
- 

## 2° The scattering Green's function

Goal: Rewrite the stationary Schrödinger equation as an integral equation that automatically incorporates the boundary condition  $\textcircled{S}$ .

The stationary Schrödinger equation reads

$$\left( -\frac{\hbar^2}{2m} \nabla^2 + V(\vec{r}) \right) \psi = E \psi, \quad E = \frac{\hbar^2}{2m} |\vec{k}|^2 > 0$$

$$\Rightarrow \underline{(\nabla^2 + k^2)} \psi = U(\vec{r}) \psi \text{ with } U = \frac{2m}{\hbar^2} V$$

We interpret this as a Helmholtz equation with a inhomogeneity  $F(\vec{r}) = U(\vec{r}) \psi(\vec{r})$ .

Def: A Green's function  $G_k(\vec{r}, \vec{r}')$  of the Helmholtz equation satisfies

$$\underline{(\nabla^2 + k^2) G_k(\vec{r}, \vec{r}') = \delta(\vec{r} - \vec{r}')}$$

Given such a Green's function, the general solution of the inhomogeneous Helmholtz eq.

$$(\nabla^2 + k^2) \psi(\vec{r}) = F(\vec{r})$$

can be constructed as

$$\psi(\vec{r}) = \psi_0(\vec{r}) + \int d^3 r' G_k(\vec{r}, \vec{r}') F(\vec{r}')$$

where  $\psi_0$  is a solution of the homogeneous equation  $(\nabla^2 + k^2) \psi_0 = 0$ .

$$\begin{aligned} \text{Proof: } (\nabla^2 + k^2) \psi &= \int d^3 r' (\nabla^2 + k^2) G_k(\vec{r}, \vec{r}') F(\vec{r}') \\ &= F(\vec{r}). \quad \square \end{aligned}$$

Claim: The Green's function can be chosen  
as (a linear combination of)

$$G_K^{(\pm)}(\vec{r}, \vec{r}') = -\frac{1}{4\pi} \frac{e^{\pm ik|\vec{r}-\vec{r}'|}}{|\vec{r}-\vec{r}'|}$$


---

Proof: Recall that  $\nabla^2\left(\frac{1}{r}\right) = -4\pi J(r)$ .

Moreover  $\nabla^2\left(\frac{1}{r}(1-e^{\pm ikr})\right) = \frac{1}{r} \frac{\partial}{\partial r} r \left[ \left( \frac{1}{r}(1-e^{\pm ikr}) \right) \right]$

$$= \frac{1}{r} k^2 e^{\pm ikr} \quad \text{because } \frac{1}{r}(1-e^{\pm ikr}) \text{ is}$$

Finite at  $r=0$ . Then

$$(\nabla^2 + k^2) \frac{1}{r} e^{\pm ikr} = (\nabla^2 + k^2) \left( \frac{1}{r} - \frac{1-e^{\pm ikr}}{r} \right)$$

$$= -4\pi J(r) - \frac{k^2}{r} e^{\pm ikr} + \frac{k^2}{r} - \frac{k^2}{r} (1-e^{\pm ikr}) =$$

$$= -4\pi J(r). \quad \square$$

---

We will show below that  $G_K^{(+)}$  and  $G_K^{(-)}$  describe outgoing and incoming spherical waves, respectively.

Using  $G_K^{(+)}$  the integral equation for  $\psi$  which corresponds to our scattering geometry reads

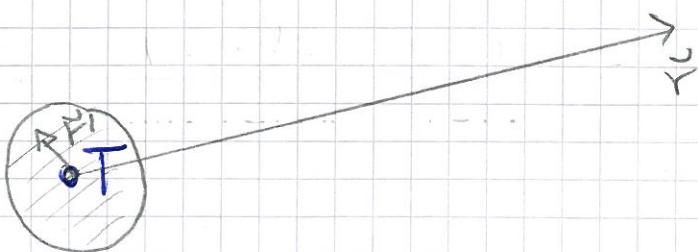
$$\underline{\psi(\vec{r})} = \underbrace{e^{ikz}}_{\text{homogeneous solution}} + \int d^3r' G_K^{(+)}(\vec{r}, \vec{r}') U(\vec{r}') \psi(\vec{r}') =$$

homogeneous solution

$$= e^{ikr} - \frac{1}{4\pi} \int d^3 r' \frac{e^{ik|\vec{r}-\vec{r}'|}}{|\vec{r}-\vec{r}'|} U(\vec{r}') \psi(\vec{r}')$$


---

To check that the solution satisfies the boundary condition (3) we note that the contributions to the integral are restricted to the target + region where  $|\vec{r}'| \ll |\vec{r}|$ :



$$\Rightarrow |\vec{r} - \vec{r}'| = \sqrt{\vec{r}^2 - 2\vec{r} \cdot \vec{r}' + \vec{r}'^2} \simeq r - \frac{\vec{r} \cdot \vec{r}'}{r}$$


---

$$\Rightarrow \int d^3 r' \frac{e^{ik|\vec{r}-\vec{r}'|}}{|\vec{r}-\vec{r}'|} U(\vec{r}') \psi(\vec{r}') \simeq$$

$$\simeq \frac{e^{ikr}}{r} \int d^3 r' e^{-ik\vec{r} \cdot \vec{r}'} U(\vec{r}') \psi(\vec{r}')$$

depends on  $\vec{r} = \frac{\vec{r}}{r} \omega(\theta, \phi)$

With this we identify

$$f(\theta, \phi) = -\frac{m}{2\pi r^2} \int d^3 r' e^{-ik_{out} \cdot \vec{r}'} V(\vec{r}') \psi(\vec{r}')$$


---

where  $k_{out} = k\vec{r}$ . But note that this still depends on the unknown solution  $\psi(\vec{r})$ .

### 3° The Born approximation

If the scattering potential is weak, the solution  $\psi(\vec{r})$  is dominated by the incoming plane wave and we may approximately write

$$\psi(\vec{r}) \approx \psi_B^{(1)}(\vec{r}) = e^{ikz} - \frac{1}{4\pi} \int d^3 r' \frac{e^{i\vec{k}(\vec{r}-\vec{r}')}}{|\vec{r}-\vec{r}'|} U(\vec{r}') e^{ikz}$$

which derives the first Born approximation. This procedure can be iterated:

$$\psi_B^{(2)}(\vec{r}) = e^{ikz} - \frac{1}{4\pi} \int d^3 r' \frac{e^{i\vec{k}(\vec{r}-\vec{r}')}}{|\vec{r}-\vec{r}'|} U(\vec{r}') \psi_B^{(1)}(\vec{r}')$$

but here we consider only the first order.

The corresponding expression for the scattering amplitude reads

$$f_B(\theta, \phi) = - \frac{1}{4\pi} \int d^3 r' e^{-i\vec{k}_{\text{out}} \cdot \vec{r}'} U(\vec{r}') e^{ikz}$$

$$= - \frac{1}{4\pi} \int d^3 r' e^{-i(\vec{k}_{\text{out}} - \vec{k}_m) \cdot \vec{r}'} U(\vec{r}')$$

where  $\vec{k}_m = k \hat{e}_z$  is the incident wave vector.

Thus  $f_B$  is proportional to the Fourier transform of the potential evaluated at

$$\vec{k} := \vec{k}_{\text{out}} - \vec{k}_m$$

$\hbar \vec{K}$  is the momentum transfer.

For radially symmetric potentials  $V(\vec{r}) = V(r)$

the integral can be evaluated by choosing  $\vec{K} = K \hat{e}_z$

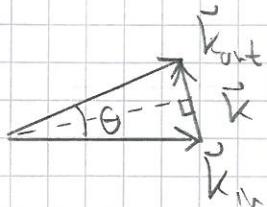
$$\int d^3r U(r) e^{-i\vec{K} \cdot \vec{r}} = \int_0^{2\pi} d\phi \int_0^\pi d\theta \int_0^\infty dr r^2 \sin\theta e^{-iKr \cos\theta} \times U(r)$$

$$= 2\pi \int_0^\infty dr r^2 U(r) \int_{-1}^1 d(\cos\theta) e^{-iKr \cos\theta}$$

$$= \frac{2}{rK} \sin(Kr)$$

$$= 4\pi K^{-1} \int_0^\infty dr r \sin(Kr) U(r)$$

$K$  is related to the scattering angle  $\theta$ :



$$\Rightarrow K = 2k \sin(\theta/2)$$

$$\Rightarrow F_B(\theta) = -\frac{1}{K} \int_0^\infty dr r \sin(Kr) U(r)$$

Example: Screened Coulomb / Yukawa potential

$$V(r) = -\frac{ze^2}{r} e^{-r/\lambda}$$

$\lambda$ : screening length

$$\Rightarrow f_B(\theta) = \frac{2m}{t^2} \frac{ze^2}{K} \int_0^\infty dr \sin(kr) e^{-r/\lambda}$$
$$= \frac{K}{K^2 + \lambda^{-2}}$$

$$= \frac{2m}{t^2} \frac{ze^2}{4k^2 \sin^2(\theta/2) + \lambda^{-2}}$$

The Coulomb case is recovered by letting  $\lambda \rightarrow \infty$

$$\Rightarrow f_B(\theta) = \frac{ze^2 m}{2(tk)^2 \sin^2(\theta/2)} = \frac{ze^2 m}{2p^2 \sin^2(\theta/2)}$$

$$\Rightarrow \frac{d\sigma}{d\Omega} = |f_B(\theta)|^2 = \frac{z^2 v^2 e^4}{4 p^4 \sin^4(\theta/2)}$$

This is the Rutherford cross section which is exact both in the classical and in the quantum mechanical scattering theory. It diverges as  $\theta \rightarrow 0$ , which implies that  $\sigma_{\text{tot}} \rightarrow \infty$ .

## Validity of the Born approximation

Replacing  $U(\vec{r})$  by  $e^{ikz}$  in the integral equation makes a difference only in the potential region, i.e. near  $\vec{r} = 0$ . We therefore demand that

$$\left| \frac{1}{4\pi} \int d^2r \frac{e^{ikz}}{|\vec{r}|} U(\vec{r}) e^{ikz} \right| \ll |e^{ikz}| = 1$$

$$= \left| \frac{1}{2} \left\{ d(\omega, \theta) \right\} \int_{-1}^1 dr \times U(r) e^{ikr} e^{ikr \omega \sin \theta} \right| =$$

$$= \frac{1}{2k} \left| \int_0^\infty dr U(r) (e^{2ikr} - 1) \right| = \frac{m}{k^2} \left| \int_0^\infty dr V(r) (e^{2ikr} - 1) \right|$$

$$\underbrace{\qquad}_{\ll 1}$$

Example: Screened Coulomb potential -  $\frac{ze^2}{r} e^{-r/\lambda}$

$$ze^2 \int_0^\infty dr \frac{1}{r} e^{-r/\lambda} (e^{2ikr} - 1) =$$

$$= \int_{\lambda k}^\infty dx e^{-xr}$$

$$= ze^2 \int_{\lambda k}^\infty dx \int_0^\infty dr e^{-xr} (e^{2ikr} - 1) =$$

$$= 2e^2 \int_{\lambda/\lambda}^{\infty} dx \left( \frac{1}{x - 2ik\lambda} - \frac{1}{x} \right) = -2e^2 \ln(1 - 2ik\lambda)$$

$$\text{and } |\ln(1 - 2ik\lambda)|^2 = \left( \frac{1}{2} \ln(1 + 4k^2\lambda^2) \right)^2 + |\operatorname{atan}(2k\lambda)|^2 \quad \}$$

We distinguish two cases:

(i) low energies,  $k\lambda \ll 1$ :

$$\ln(1 + 4k^2\lambda^2) \approx 4k^2\lambda^2, \operatorname{atan}(2k\lambda) \approx 2k\lambda$$

$$\Rightarrow \frac{w}{\hbar^2 k} \cdot 2e^2 |2k\lambda| = \frac{2e^2}{\lambda} \cdot \frac{2w\lambda^2}{\hbar^2} \ll 1$$

Interpretation: The typical potential energy at  $r \approx \lambda$  is

$$V_0 \approx \frac{2e^2}{\lambda}$$

and the kinetic energy of a particle confined to that scale is

$$E_{kin} \approx \frac{\hbar^2}{2m\lambda^2}$$

$$\Rightarrow V_0 / E_{kin} \ll 1$$

which means that the potential is too weak to form a bound state (in three dimensions)

(ii) high energies,  $k\lambda \gg 1$ :

$$|Q(1-2ik\lambda)| \approx \frac{1}{2} \ln(4k^2\lambda^2)$$

which increases very slowly with  $k\lambda$ . Thus the condition reads essentially

$$\frac{mv}{k^2\lambda} \cdot 2e^2 = 2 \left( \frac{e^2}{Ec} \right) \frac{mc}{\hbar k} = 2\alpha \frac{c}{v} \ll 1$$

where  $v = p/m$  is the velocity of the incident particle and  $\alpha$  is the fine structure constant.

Thus in this case the approximation is valid for light nuclei and high velocities.

#### 4° Partial wave analysis

Goal: Represent the scattering problem for spherically symmetric potentials in terms of spherical harmonics.

#### a) Radial Schrödinger equation for free particles

Premise: The Schrödinger Hamiltonian in spherical coordinates reads

$$H = -\frac{\hbar^2}{2m} \frac{1}{r} \frac{\partial^2}{\partial r^2} r + \frac{L^2}{2mr^2} + V(r)$$